

Indirect Measurements of the $^{18}\text{F}(\alpha, p)^{21}\text{Ne}$ Reaction with the TUDA Scattering Chamber

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Astrophysics with modern small-scale accelerators



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Astrophysical Motivation

Impact on ^{19}F abundances in AGB models

- Region of interest is the He burning shell in thermally pulsing AGB stars.
- AGB models do not synthesise enough ^{19}F to match F/O abundances observed.

How can the $^{18}\text{F}(\alpha, p)^{21}\text{Ne}$ reaction help?

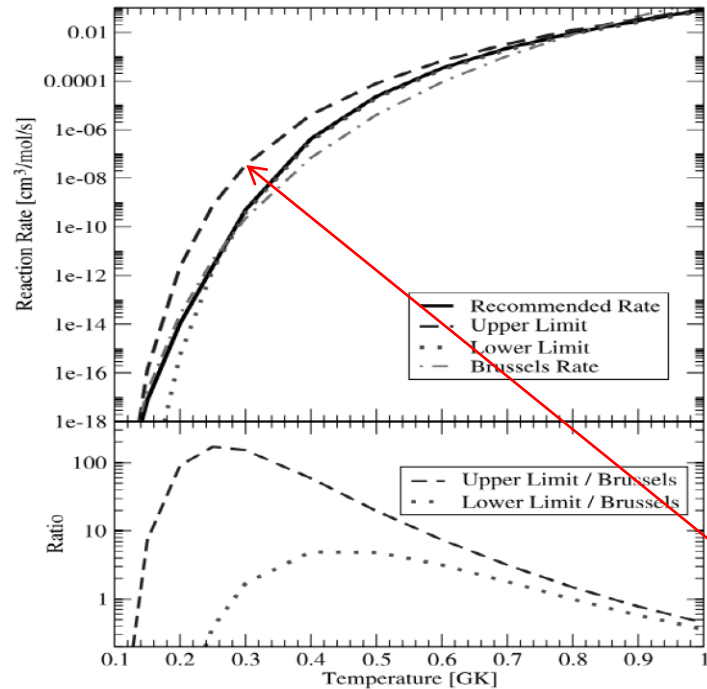
- $^{18}\text{F}(\alpha, p)^{21}\text{Ne}$ reaction competes with $^{18}\text{F}(\beta^+ \nu)^{18}\text{O}$ as well as producing protons.
- This increases the product $N_{180}N_p$
- Enhances production of ^{19}F via the chain $^{18}\text{O}(p, \alpha)^{15}\text{N}(\alpha, \gamma)^{19}\text{F}$



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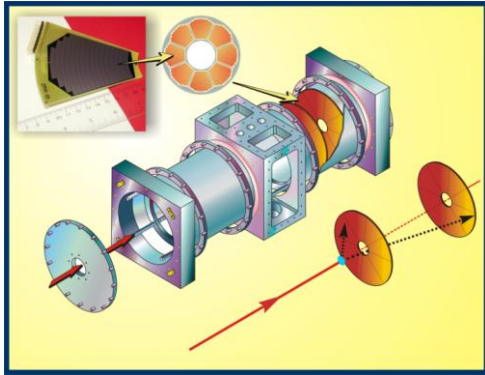
Previous Measurements



Amanda I. Karakas *et. al.* 2008

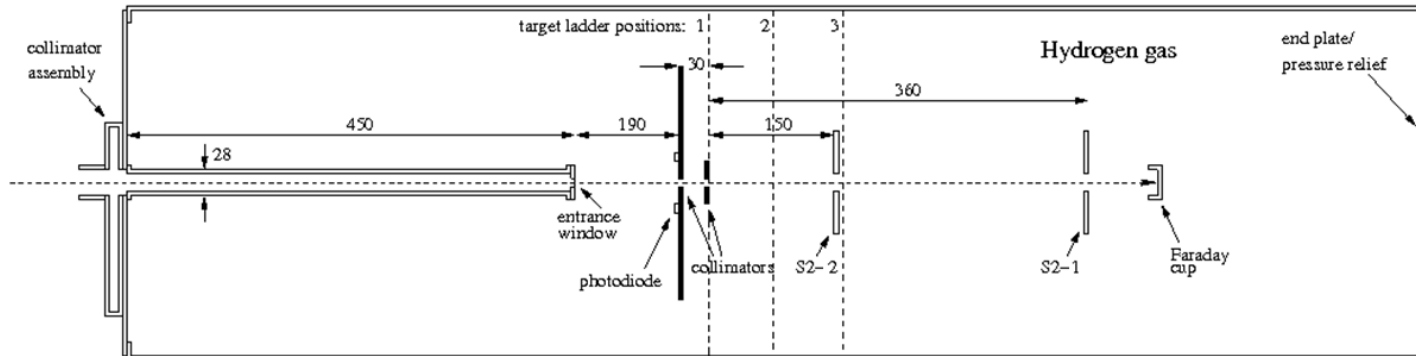
- Time-reversed measurement measured by H. Y. Lee *et. al.* using proton beam and ²¹Ne implanted Cu/Au targets.
- Models suggest rate lies closer to upper limit.
- Systematic error in this data from oxygen contamination on backing material via:
 $^{17}\text{O}(p,\gamma)^{18}\text{F}$
 $^{18}\text{O}(p,n)^{18}\text{F}$
- This gives a high upper limit at energies of interest.

Inverse reaction measurement with H₂ target



<http://tuda.triumf.ca/>

- ²¹Ne beam on hydrogen gas target to measure ²¹Ne(p,α)¹⁹F.
- Safety procedures implemented.
- Beam energies ranging from 50.6 to 68.2 MeV corresponding to energies of 0.6 to 1.4 MeV in the ¹⁸F+α system.



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Thank you!



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